Doppler Effect and Working With Redshifts

Name

Purpose:

 To provide you with necessary skills to understand how to calculate redshifts and apply the Doppler Effect

Estimated Completion Time: 45 minutes

Resources needed:

- Calculator (preferably scientific)
- Textbook
- Web access is highly desirable

Questions

1. Complete the following table (5 marks):

Galax	Spectral Line	Rest λ	Observed λ	Z	Approaching	Velocity (km/s)
У	Observed	(nm)	(nm)		or receding?	
1	Ηα	656.3	670			
2	H (calcium)	396.9	380			
3	K (calcium)	393.4	410			
4	Ηβ	486.1	600			
5	Hγ	434.1	800			

- 2. For which galaxies in question 1 was it necessary to use the relativistic correction? See <u>http://www.kcvs.ca/martin/astro/kingsu/unit5/141/z_rel.html</u>. (2 marks)
- 3. The H line of calcium is a commonly used spectral line in galaxies. Suppose you measure the spectrum of a galaxy and find that $\lambda = 450$ for the H line. Is this galaxy moving toward you or away from you and at what speed? (3 marks)

1/2

4. The following figure shows arrows that represent the rotational velocity of a spiral galaxy measured at various locations across the galaxy. Sketch what the rotational velocity graph for this galaxy would look like and explain why it is shaped the way it is. Note the direction and length of the arrows. Also – your answer will just be qualitative – you are not expected to provide actual numbers on your graph. You may wish to consult the on line lectures

http://www.kcvs.ca/martin/astro/kingsu/unit5/132/chp13_2.html (4 marks)



Other useful links:

http://www.kcvs.ca/martin/astro/kingsu/unit2/63/chp6_3.html http://www.kcvs.ca/martin/astro/kingsu/unit5/132/chp13_2.html